

Lunar Reflective Calibration Beacon for Radio Astronomy and SETI

Project: EME PI: Dr. H. Paul Shuch, Exec. Dir., The SETI League, Inc. (paul@setileague.org)

Description and Objectives:

The SETI League EME Beacon consists of a microwave transmitter with antennas automatically tracking the Moon across the sky, reflecting to Earth a precision weak signal to calibrate radio telescopes worldwide.

Key Features:

- High frequency precision, locked to atomic clock
- Transmission in the 23 cm amateur radio band
- Automatic az-el antenna tracking of Moon
- Remotely programmable modulation and power
- Continuous, unattended operation

Approach:

Location: Kinnelon NJ USA, Grid Square FN21ta Station Trustee: Richard Factor, WA2IKL Transmitter: currently 350 watts CW output (nominal) Frequency: 1296.000 MHz Accuracy: better than +- 2 Hz Feedline: 60 feet of 5/8 inch Cellwave hardline (estimated loss 3 dB) Antenna: quad helix array, RHCP uplink; gain +24 dBi Tracking Hardware: Kansas City Tracker driving Yaesu az/el rotors Tracking Sortware: NOVA for Windows EIRP: estimated at +76 dBm

<u>Partners:</u>

- American Astronomical Society
- ARRL Foundation
- SETI Institute



Schedule Milestones and Deliverables:	Target:
04 Mar 2001: First Light, 20 Watt IPA	TRL 4
09 Mar 2001: Successful Arecibo tests	TRL 5
17 Feb 2002: First Light, 200 Watt PA	TRL 6
13 Apr 2003: Arecibo/Jodrell Bank FUDD tests	TRL 7
15 Nov 2003: Released for general use worldwide	TRL 8
13 Mar 2006: Upgraded, 350 Watt PA	TRL 9

Applications:

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- Time and Frequency Standard for Project Argus
- Sensitivity calibration for amateur radio telescopes
- End-to-end system verification for Project Phoenix
- Sensitivity calibration for professional observatories
 - Revised: 13 March 2006

Keywords: EME, Moonbounce, Microwave Beacon, Transmitter, Radio Astronomy, SETI, L-Band, Arecibo

TRL